

VI Workshop Challenges in New Physics in Space

From Quantum to Classical Instability in Relativistic Stars

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In collaboration with W. Lima, S. Mastreas, and D. Vanzella

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AWAKING THE VACUUM

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❖ It has been shown that relativistic stars may become unstable due to quantum effects: The vacuum awaking effect [W. Lima, D. Vanzella, Phys. Rev. Lett. 104, 161102 (2010), W. Lima, G. Matsas, D. Vanzella, Phys. Rev. Lett. 105, 151102 (2010), W. Lima, R. Mendes, G. Matsas, D. Vanzella, Phys. Rev. D 87, 104039 (2013)]

AWAKING THE VACUUM

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$$\Sigma_{t_1}$$



$$ds^2 = -dt^2 + dx^2$$

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Σ_{t_2}



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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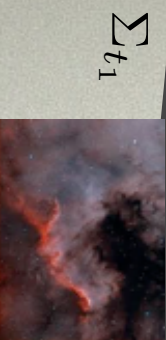
AWAKING THE VACUUM



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$$\langle \hat{\Phi}^2 \rangle_{\rm in} \sim \frac{\kappa e^{2\bar{\Omega}t}}{2\bar{\Omega}} \left(\frac{\bar{F}(\mathbf{x})}{f_{\rm int}(\mathbf{x})} \right)^2 [1 + \mathcal{O}(e^{-\epsilon t})]$$



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$$\langle \hat{\rho} \rangle_{\text{in}} \sim \kappa \frac{\bar{\Omega} e^{2\bar{\Omega}t}}{16\pi\sqrt{f}} \left\{ \frac{1-4\xi}{2r^2} \frac{d}{d\chi} \left(r^2 \frac{d}{d\chi} \left(\frac{F_{\bar{\Omega}0}}{\bar{\Omega}r} \right)^2 \right) + \frac{\xi}{\bar{\Omega}^2 r^2} \frac{d}{d\chi} \left(\frac{F_{\bar{\Omega}0}}{f} \frac{df}{d\chi} \right) \right\} [1 + \mathcal{O}(e^{-\epsilon t})].$$

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- ❖ The unstable phase must be detained by backreaction effects which should be responsible to bring the vacuum back to some stationary regime.

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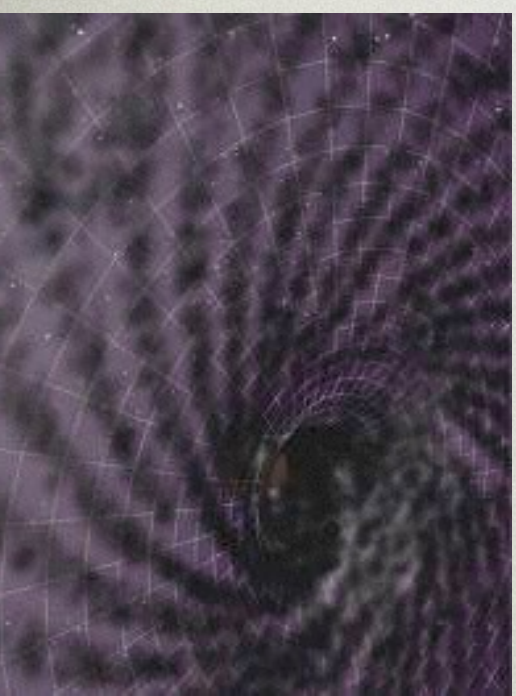
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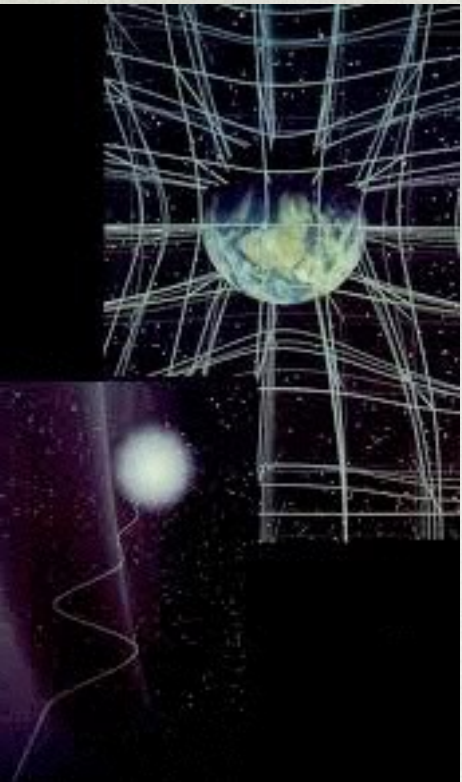


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- ♣ Einstein semiclassical equation:

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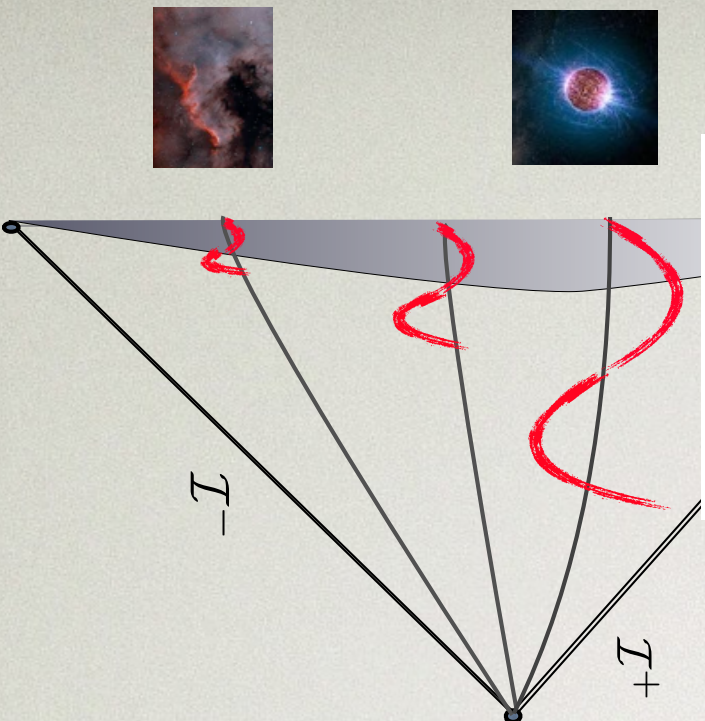
This approach, however, have many difficulties, e.g.,

(1) One does not expect it to be valid when the fluctuations of the stress energy tensor are “large” when compared to its mean value

(2) Even if/when it is valid, it is, in general, prohibitively difficult to solve it.

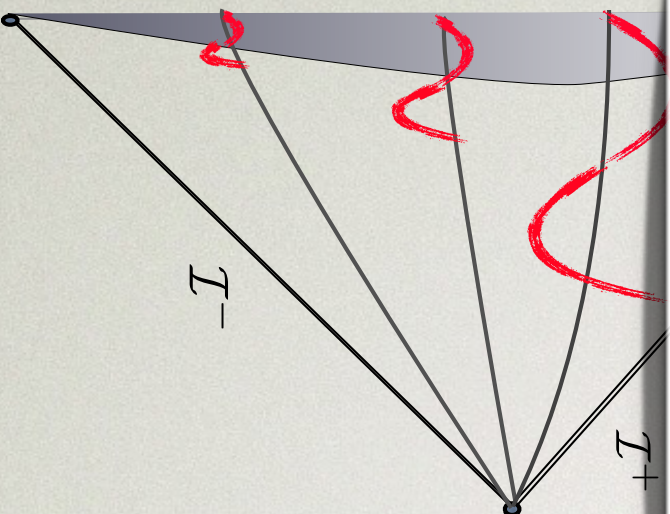
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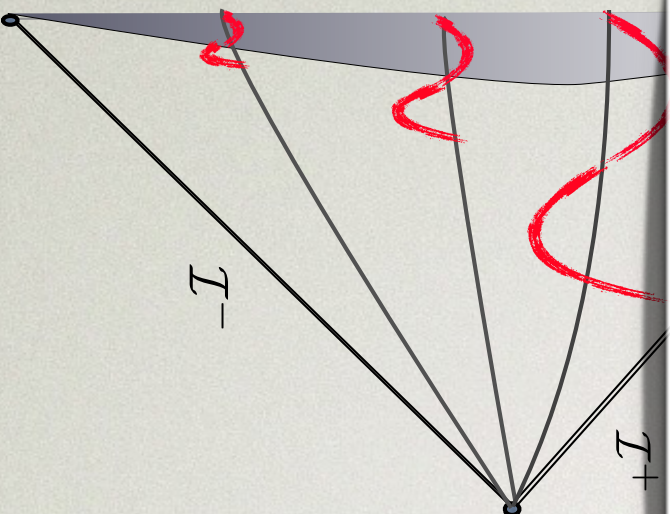
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it seems reasonable to believe that quantum fluctuations amplified enough to menace the stability of relativistic stars cannot remain "quantum" for too long.



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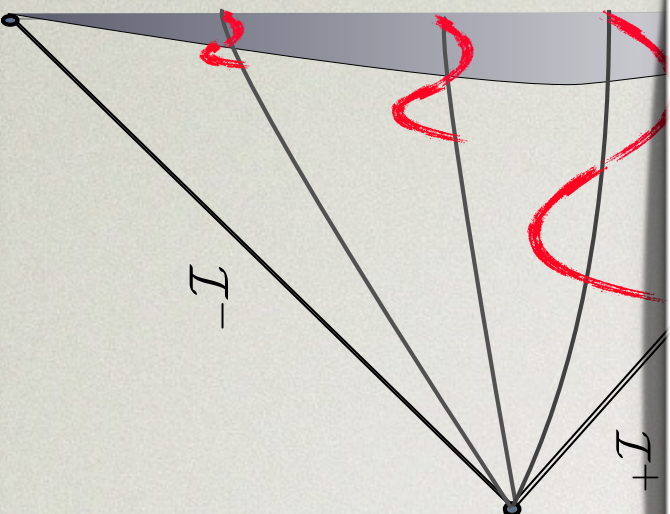
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If the quantum phase ends before vacuum fluctuations dominate the system, we expect backreaction to be well described by the classical general-relativistic equations.

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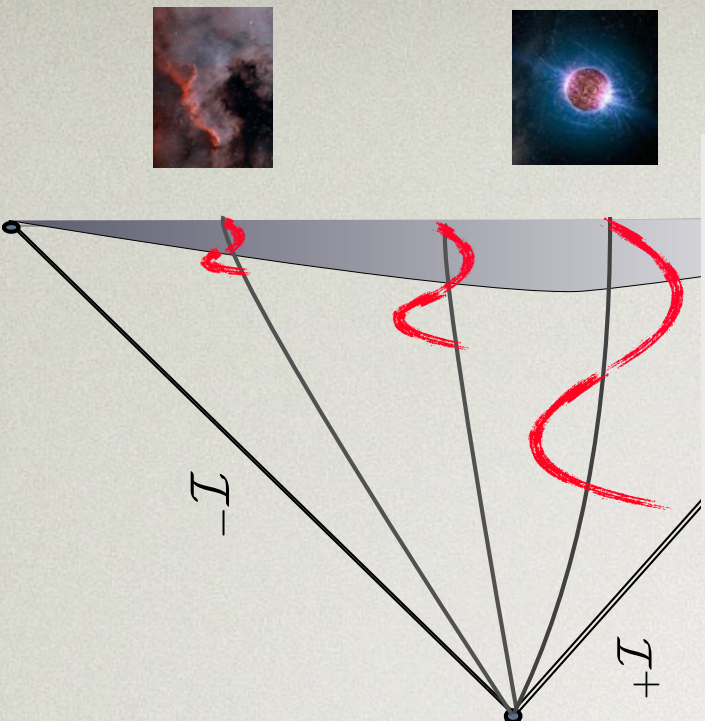
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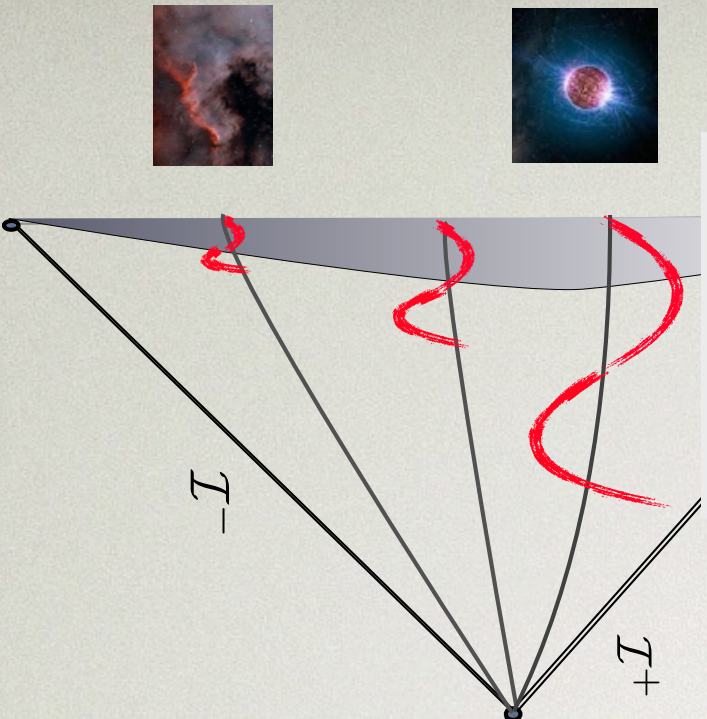
The relationship between the quantum and classical description of the instability was analyzed by R. Mendes, G. Matsas, and D. Vanzella PRD (1014)

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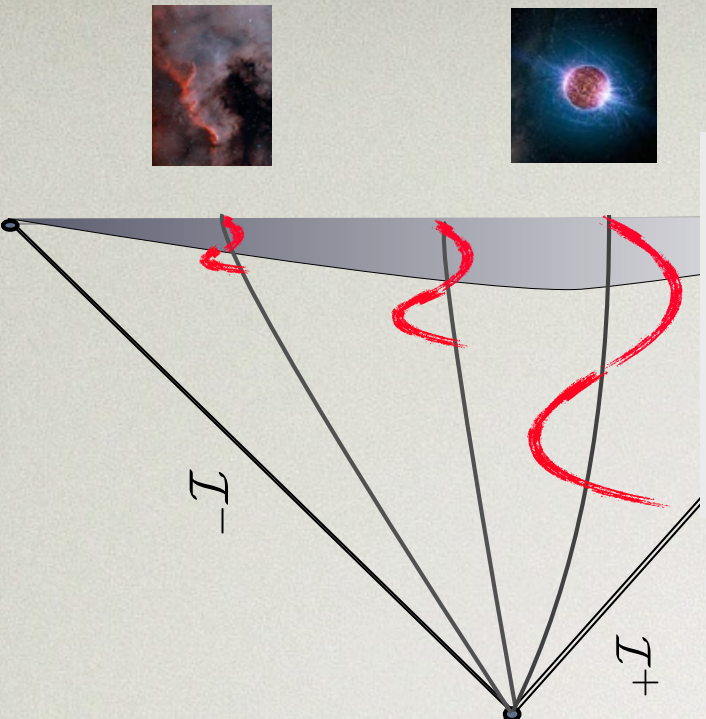
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(1) The appearance of certain "classical" correlations

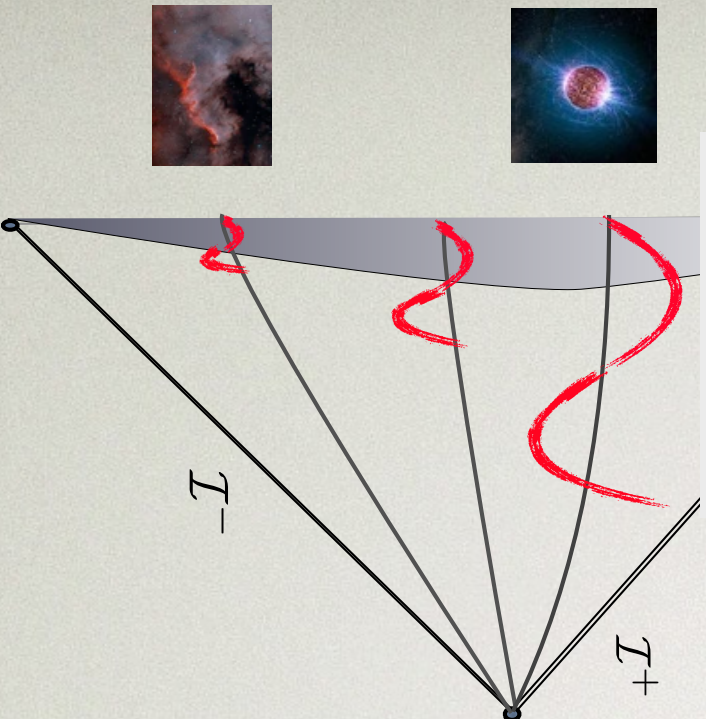


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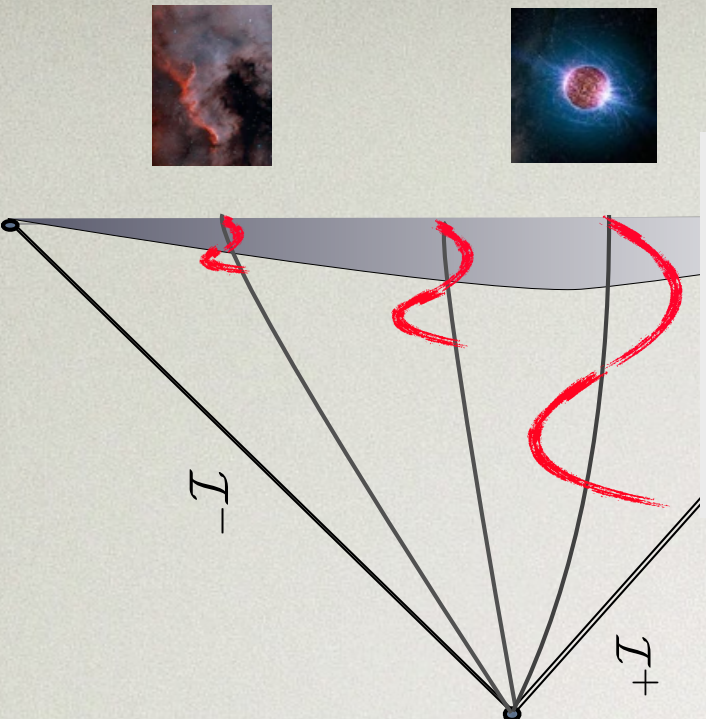
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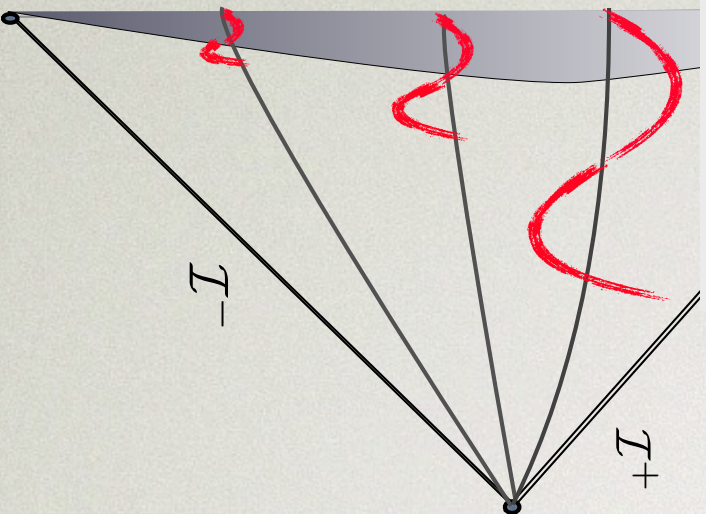
System
 S



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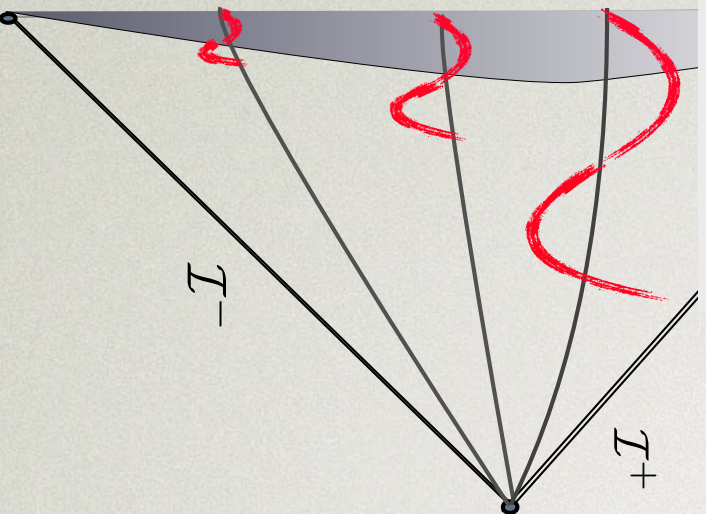
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System S + Environment $\mathcal{E}$

AWAKING THE VACUUM

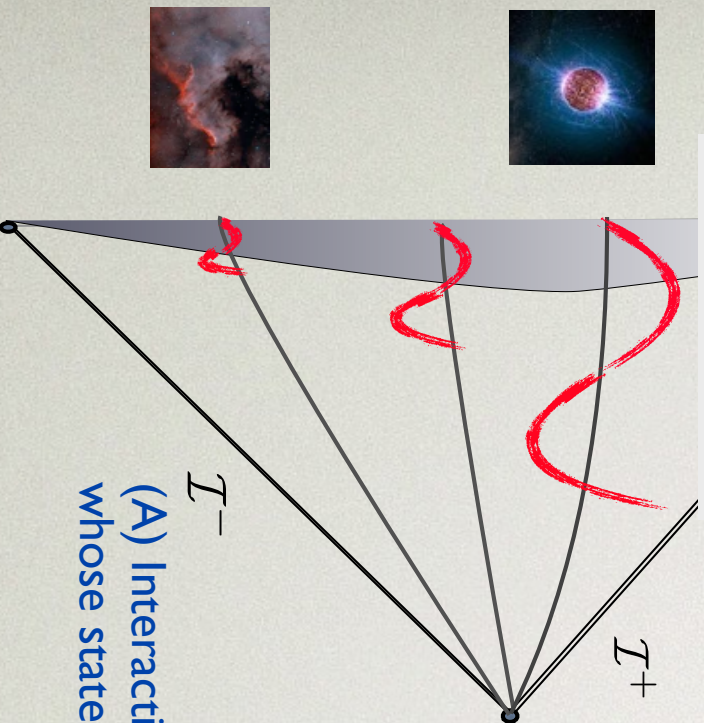
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(A) Interaction between $\mathcal{S}$ and $\mathcal{E}$ selects a preferred basis of $\mathcal{S}$ whose states are “stable” (in spite of the interaction with $\mathcal{E}$)



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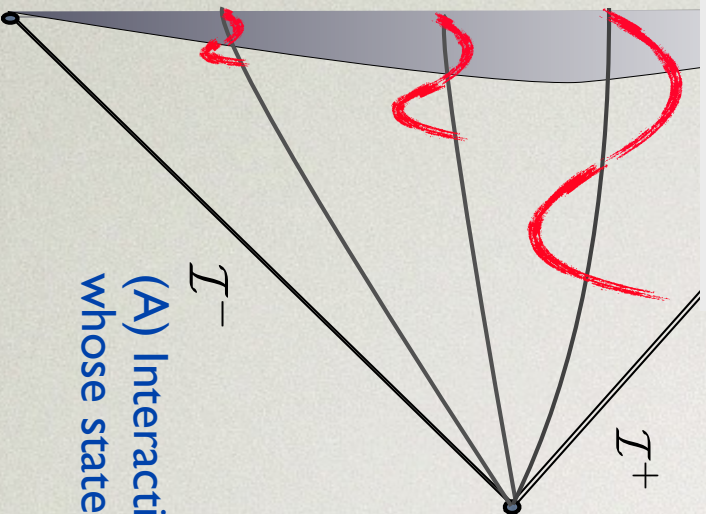
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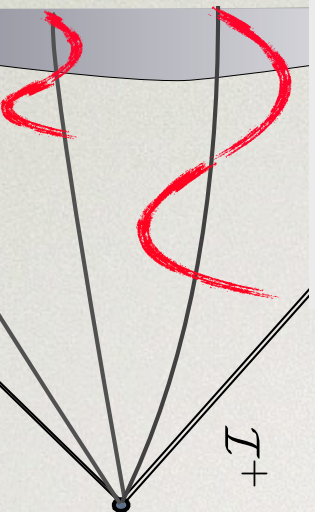
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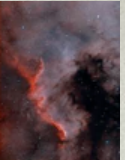
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System $\mathcal{S}$ + Environment $\mathcal{E}$

T^-



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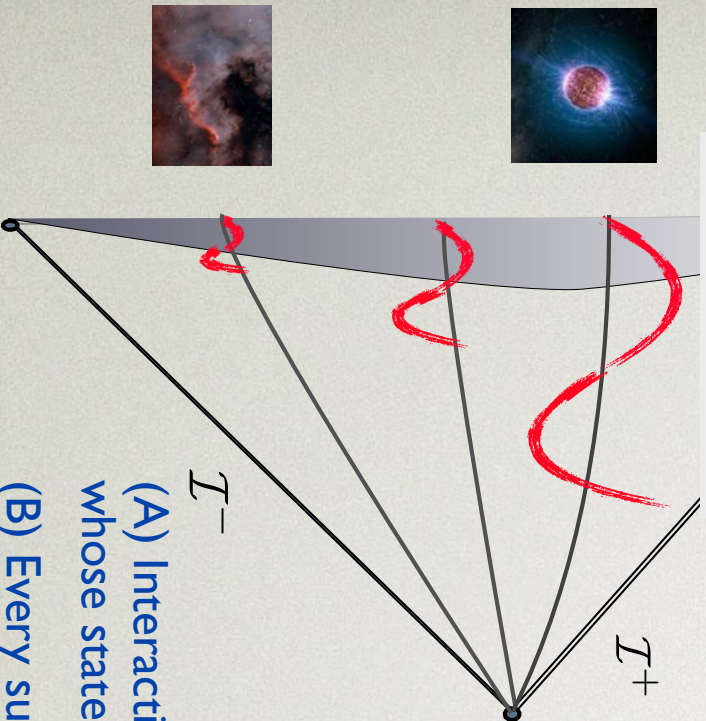
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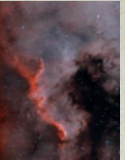
(B) Every superposition of pointer states evolves to a statistical mixture of the pointer states (i.e. every density matrix become diagonal in the pointer state basis)



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$\mathcal{I}^-$

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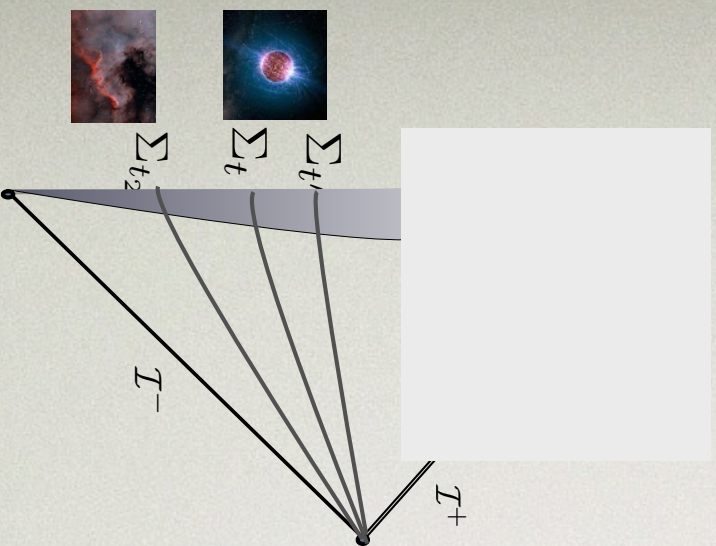
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W. Zurek, RMP (2003), W. Zurek, Nature Phys. (2009)

FROM QUANTUM TO CLASSICAL

A. Landolfo, W. Lima, S. Mastias, and D. Vanzella, PRD (2015)

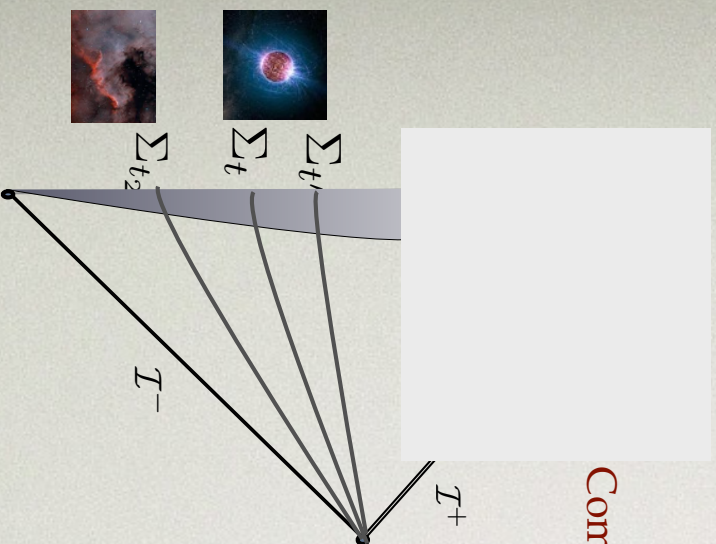


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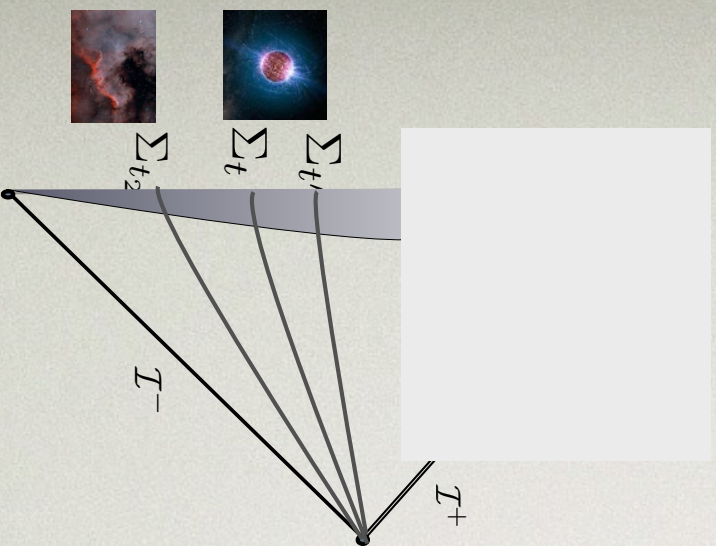
Complete orthonormal set of solutions
of KG equation

$$\{v_{\omega l \mu}^{(+)}, v_{\omega l \mu}^{(-)}\}, \{w_{\Omega}^{(+)}, w_{\Omega}^{(-)}\}$$



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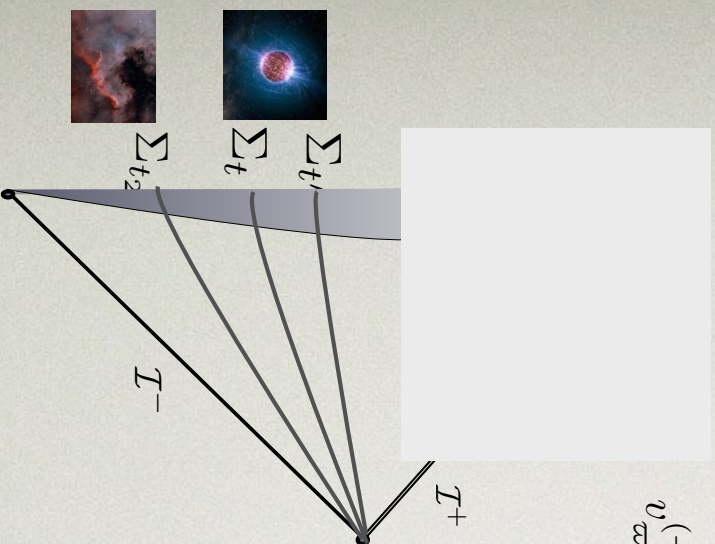
A. Jorjadze, W. Jma, S. Mastas, and D. Panzella, PRD (2015)



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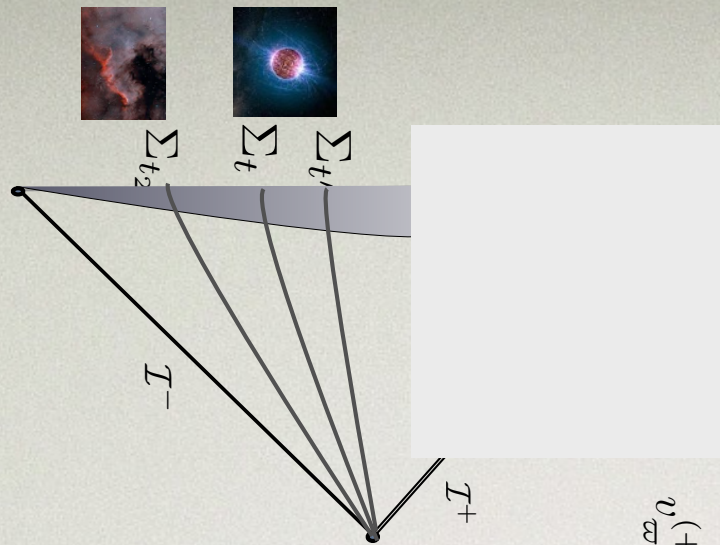
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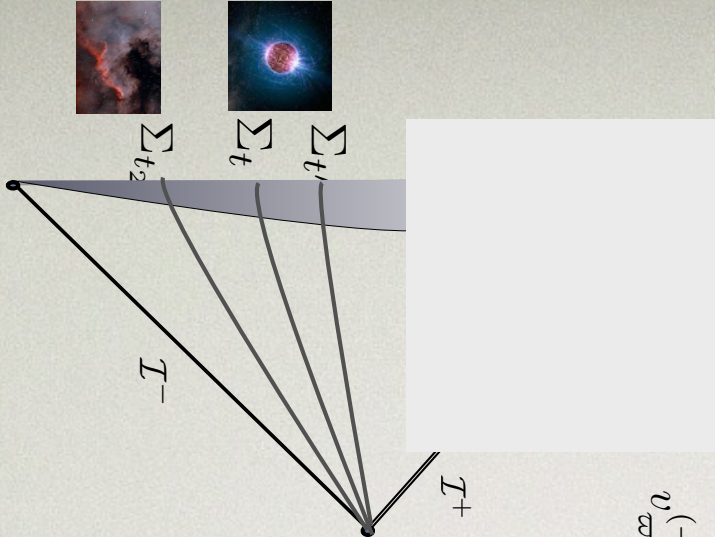


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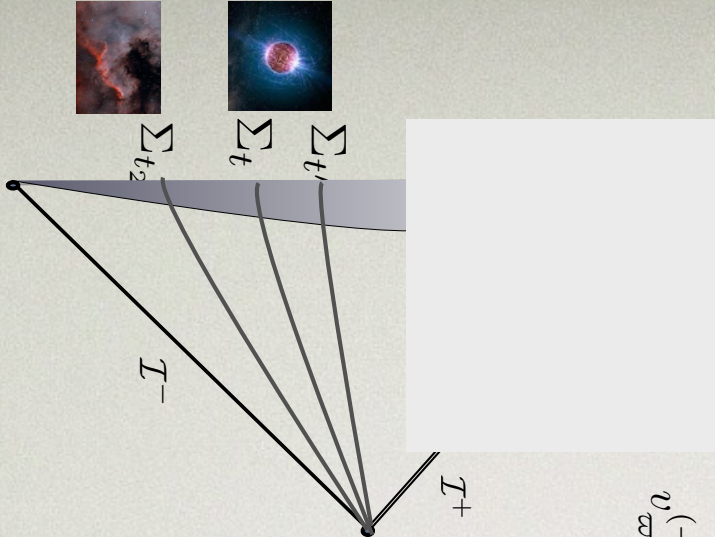
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Hamiltonian:

$$\hat{H} = \hat{H}_s + \hat{H}_u$$

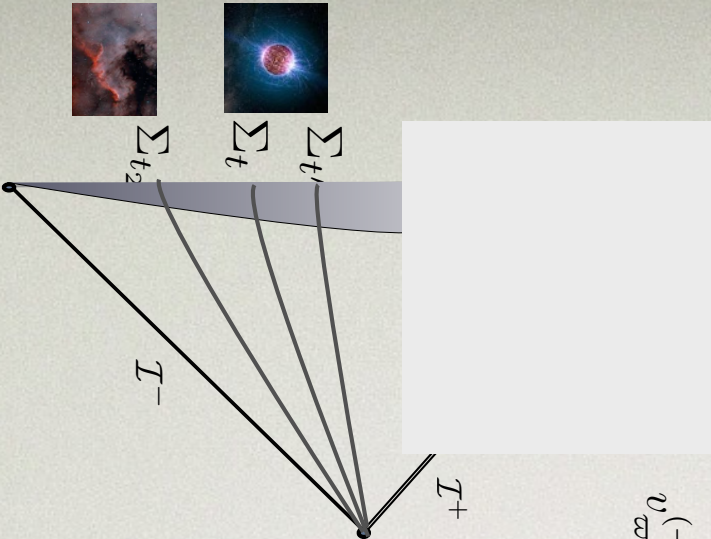


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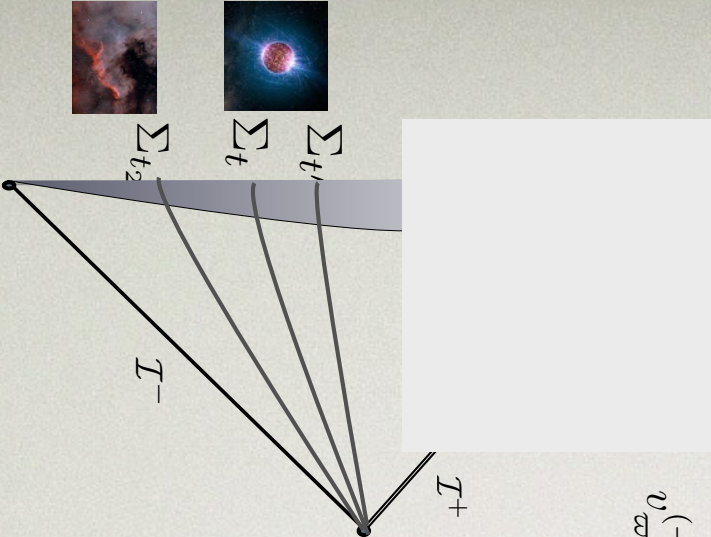
$$\hat{H}_s = \frac{1}{2} \sum_{l\mu} \int d\varpi [\hat{b}_{\varpi l\mu} \hat{b}_{\varpi l\mu}^{\dagger} + \hat{b}_{\varpi l\mu}^{\dagger} \hat{b}_{\varpi l\mu}] \varpi$$

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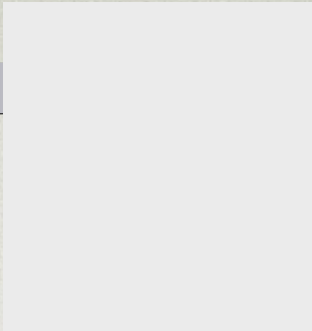
$$\hat{H}_u=-\frac{\Omega}{2}(\hat{a}_{\Omega}\hat{a}_{\Omega}+\hat{a}_{\Omega}^{\dagger}\hat{a}_{\Omega}^{\dagger})$$

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$$\hat{H}_u = -\frac{\Omega}{2} (\hat{a}_{\Omega} \hat{a}_{\Omega} + \hat{a}_{\Omega}^{\dagger} \hat{a}_{\Omega}^{\dagger})$$

$$\hat{q}_{\Omega} \equiv \frac{1}{\sqrt{2\Omega}} (\hat{a}_{\Omega} + \hat{a}_{\Omega}^{\dagger})$$

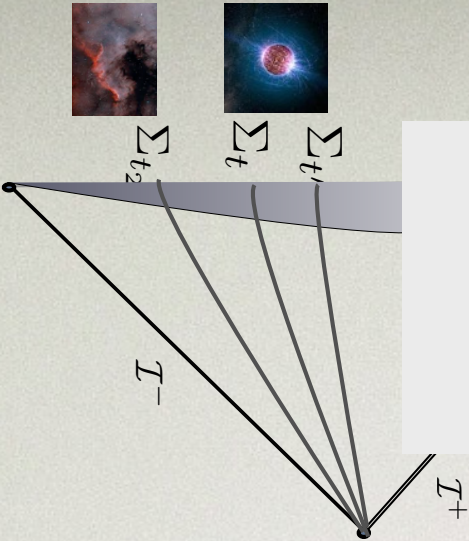


FROM QUANTUM TO

CLASSICAL

A. Landulfo, W. Lima, S. Mastasas, and D. Vanzella, PRD (2015)

$$v_{\varpi l\mu}^{(+)} = e^{-i\varpi t} \frac{F_{\varpi l}(\chi)}{\sqrt{2\varpi} r(\chi)} Y_{l\mu}(\theta, \phi), \quad w_{\Omega}^{(+)} = \frac{e^{\Omega t - i\pi/4} + e^{-\Omega t + i\pi/4}}{\sqrt{4\Omega}} \frac{F_{\Omega}(\chi)}{r(\chi)} Y_{00}(\theta, \varphi)$$



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Hamiltonian:

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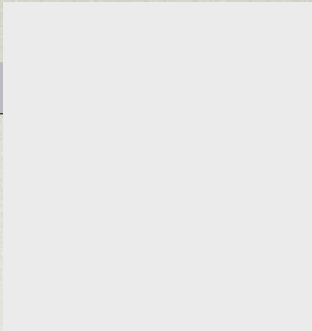
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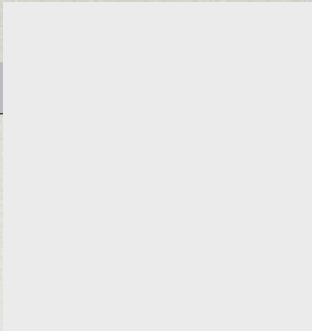


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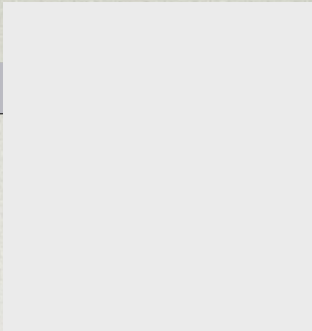


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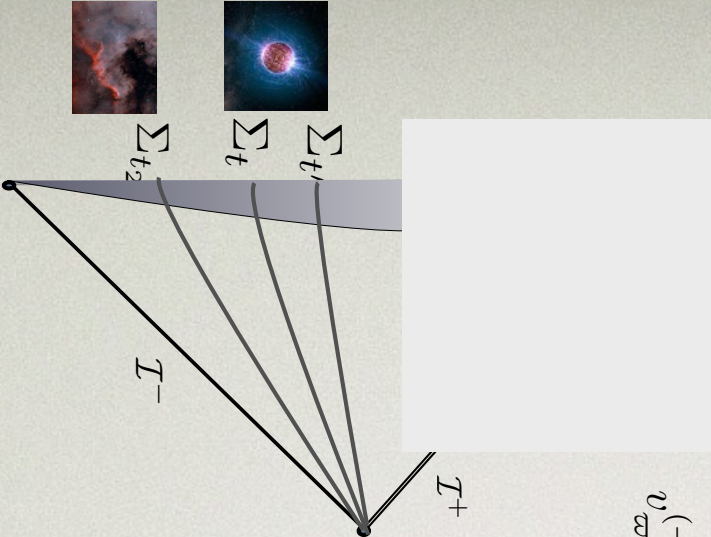
Upside down harmonic oscillator

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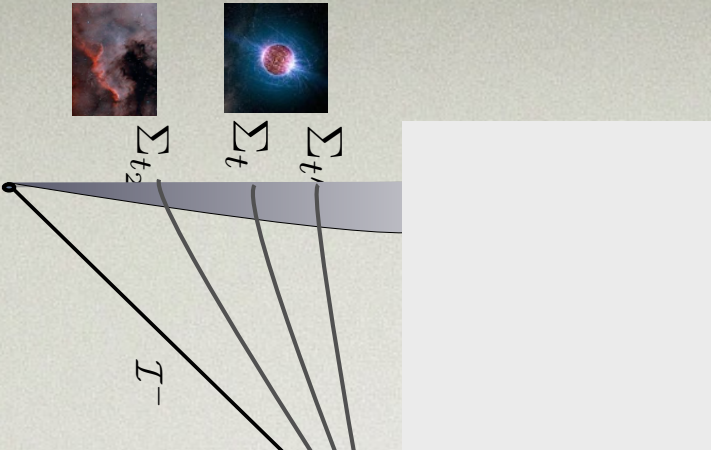
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FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

In this fixed background, the modes decouple and evolve independently.

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase



Σ_t

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

❖ Field Fock Space

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FROM QUANTUM TO CLASSICAL

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Σ_t

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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$$\mathfrak{F}_s(\mathcal{H}_s) \otimes \mathfrak{F}_s(\mathcal{H}_u)$$

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FROM QUANTUM TO CLASSICAL

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Σ_t



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- ❖ Unstable Phase Vacuum

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FROM QUANTUM TO CLASSICAL

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$$\hat{U}(t) \equiv e^{-it\hat{H}} = e^{-it\hat{H}_s} \otimes e^{-it\hat{H}_u}$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

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FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



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Squeezed Vacuum

$$U_u(t) \equiv e^{-it\hat{H}_u}$$

Squeezing Operator

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



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FROM QUANTUM TO CLASSICAL

❖ Unstable Phase



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$$a(t) \equiv \frac{\Omega}{\cosh 2\Omega t}$$

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FROM QUANTUM TO CLASSICAL

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FROM QUANTUM TO CLASSICAL

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Σ_t

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

- ❖ Wigner Function

$$W(t, q, p) \equiv \frac{1}{2\pi} \int_{-\infty}^{+\infty} dy \varrho(t, q - y/2, q + y/2) e^{ipy},$$

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- ❖ For quadratic potentials (like our upside down harmonic oscillator)

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It is a very useful tool to analyze the classicalization of a quantum system

Liouville equation

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



- ❖ Wigner Function

$$W(t, q, p) \equiv \frac{1}{2\pi} \int_{-\infty}^{+\infty} dy \varrho(t, q - y/2, q + y/2)$$

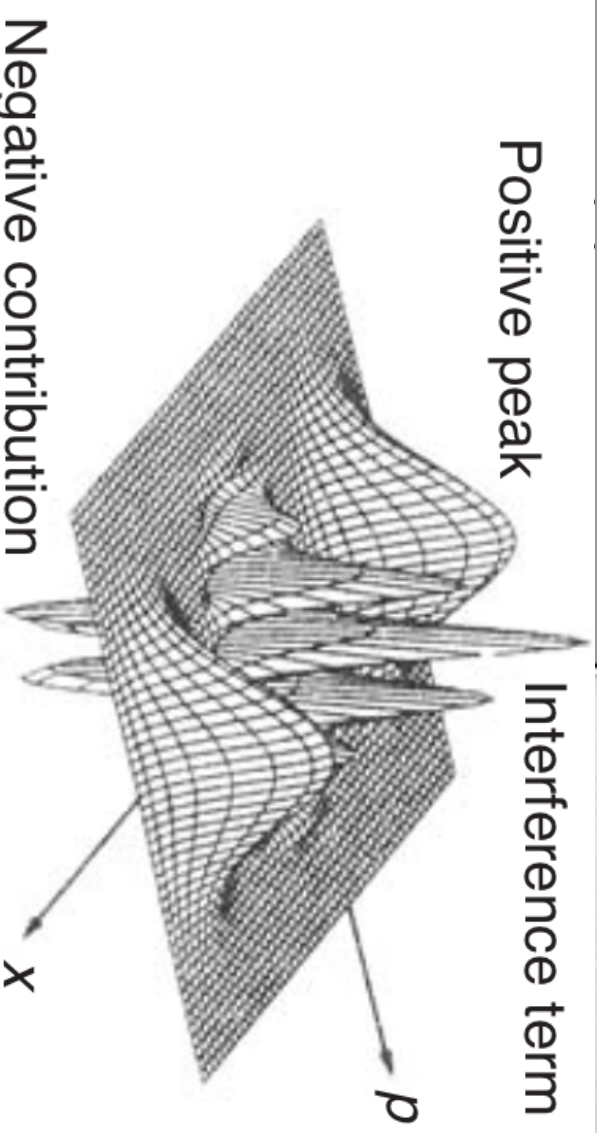
- ❖ For pure states $|\psi\rangle$

$$\int_{-\infty}^{+\infty} dq W(q, p) = |\tilde{\psi}(p)|^2, \quad \int_{-\infty}^{+\infty} dp W(q, p) = |\psi(q)|^2$$

- ❖ For quadratic potentials (like our ups

$$\partial_t W(t, q, p) = \{H(q, p), W(t, q, p)\}$$

It is a very classical



Negative contribution

Positive peak

Interference term

W. Zurek, Decoherence and the Transition from Quantum to Classical—Revisited

Liouville equation

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



- ❖ For the upside down harmonic oscillator

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



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- ❖ For the upside down harmonic oscillator

$$H(q, p) \equiv \frac{1}{2}p^2 - \frac{\Omega^2}{2}q^2.$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase



Σ_t

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FROM QUANTUM TO CLASSICAL

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- ❖ Squeezed vacuum

FROM QUANTUM TO CLASSICAL

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$$\eta(t,q) \equiv \langle q | \eta(t) \rangle = \left(\frac{a(t)}{\pi} \right)^{1/4} e^{i\alpha(t)q^2/2}.$$

FROM QUANTUM TO CLASSICAL

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FROM QUANTUM TO CLASSICAL

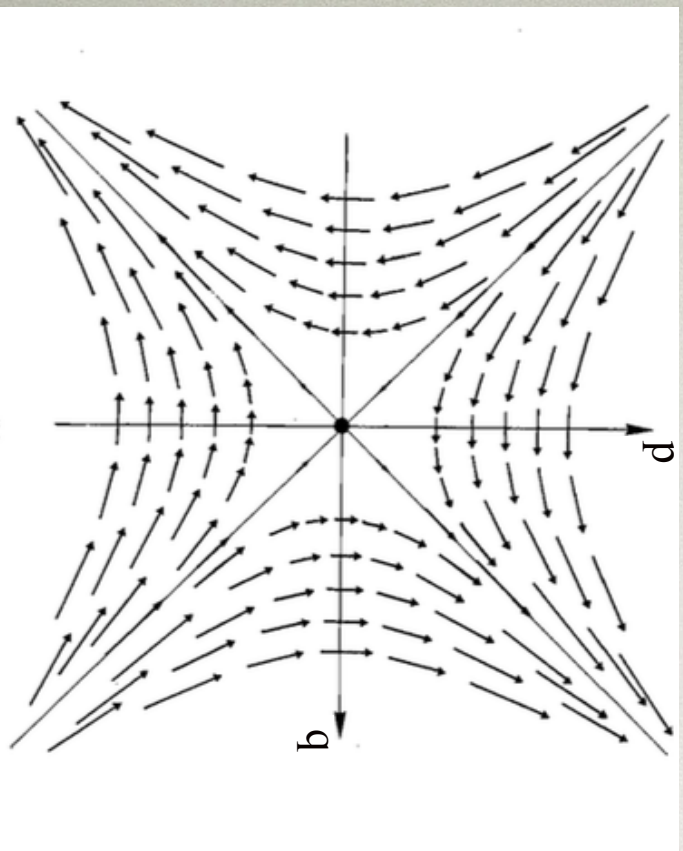
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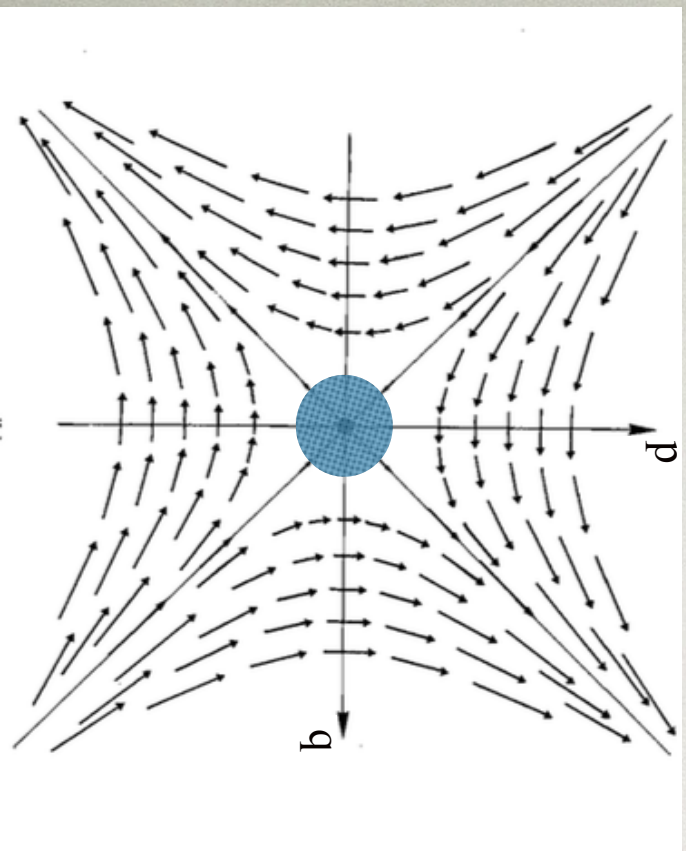
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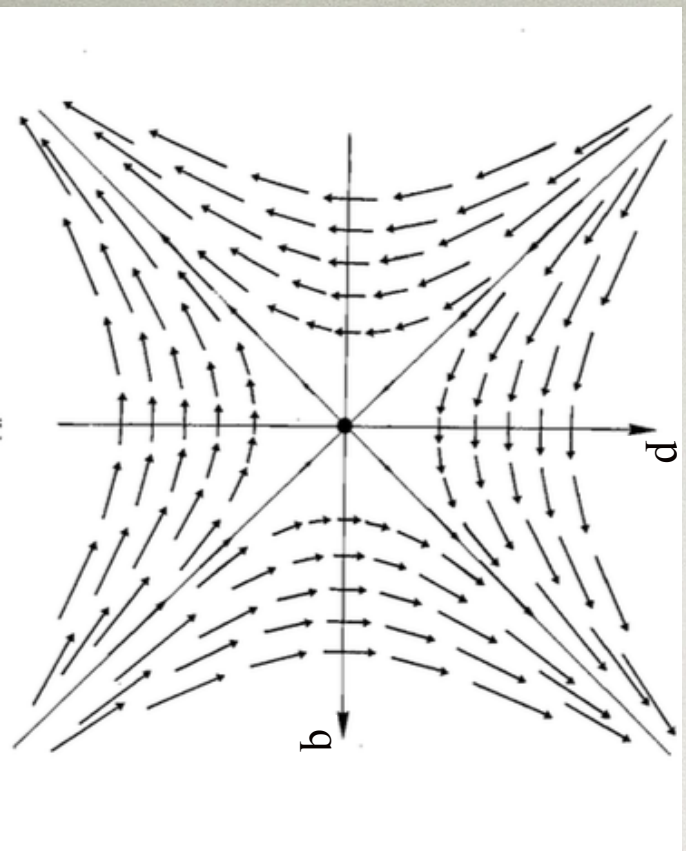
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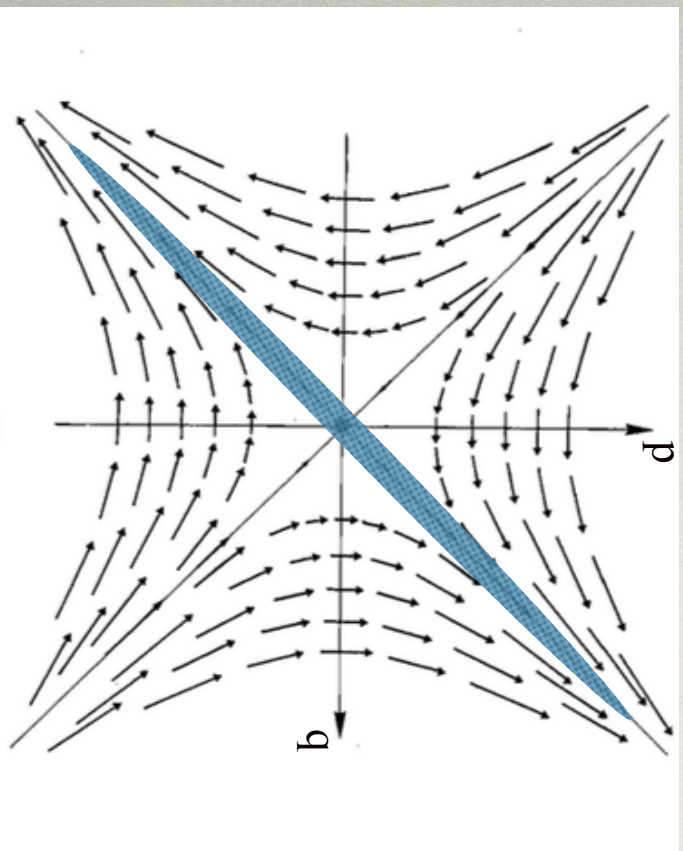
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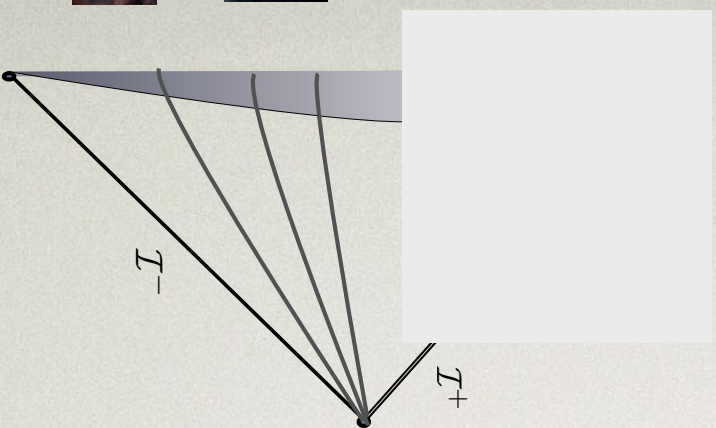
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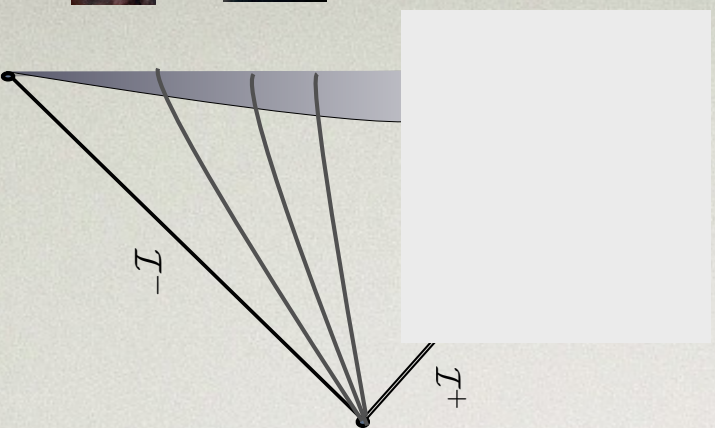
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FROM QUANTUM TO CLASSICAL



FROM QUANTUM TO CLASSICAL

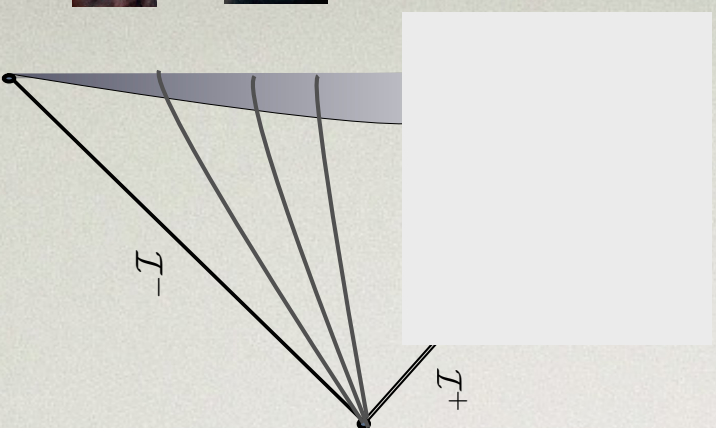
$$S = S_{\text{EH}} + S_{\Phi} + S_{\text{M}}$$



FROM QUANTUM TO CLASSICAL

$$S = S_{\text{EH}} + S_{\Phi} + S_{\text{M}}$$

$$S_{\text{EH}}[g_{ab}] \equiv \frac{2}{\kappa^2} \int_{\mathcal{M}} d^4x \sqrt{-g} R, \quad \kappa \equiv \sqrt{32\pi G}$$

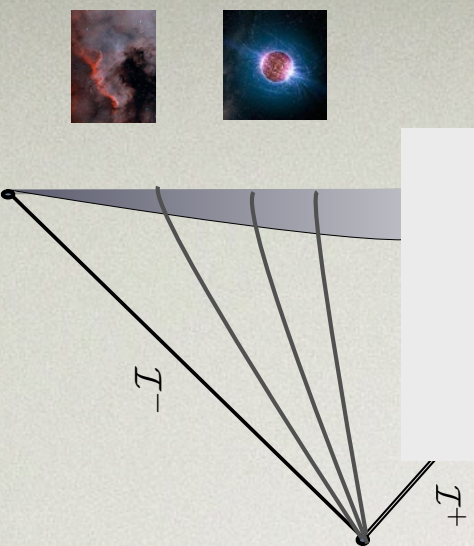


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Einstein-Hilbert action

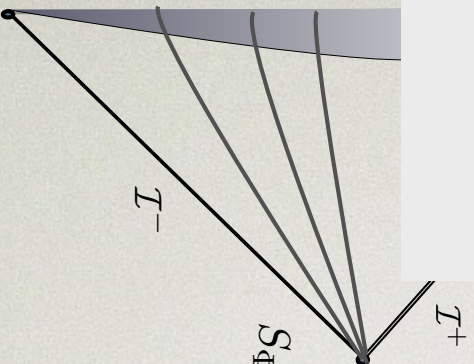


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Einstein-Hilbert action

$$S_{\Phi}[\Phi, g_{ab}] \equiv -\frac{1}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} [\nabla_a \Phi \nabla^a \Phi + (m^2 + \xi R) \Phi^2],$$



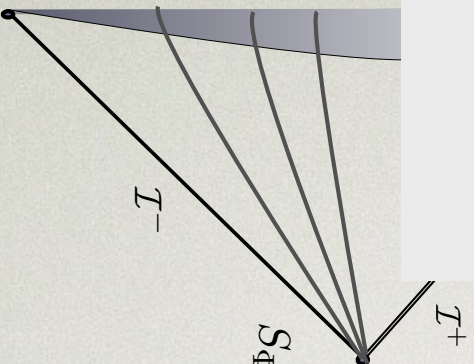
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Scalar field action



FROM QUANTUM TO CLASSICAL

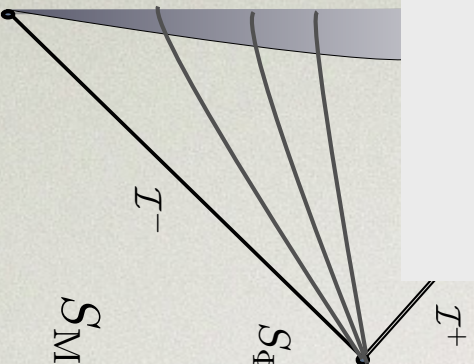
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Scalar field action

$$S_{\text{M}}[\Psi, g_{ab}]$$



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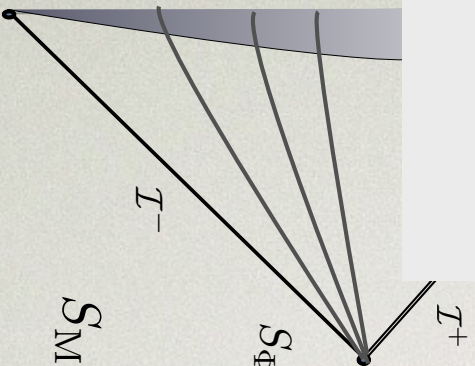
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scalar field action

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classical-matter action



FROM QUANTUM TO CLASSICAL

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Einstein-Hilbert action

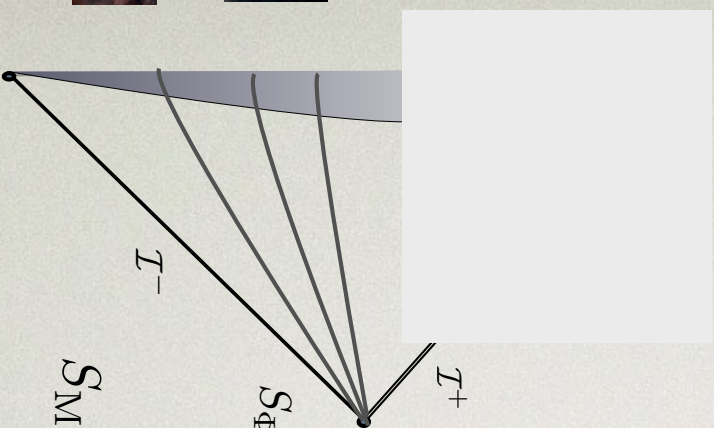
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scalar field action

$$S_{\text{M}}[\Psi, g_{ab}]$$

classical-matter action

$$g_{ab} \rightarrow g_{ab} + \kappa \gamma_{ab}, \quad \Phi \rightarrow \Phi + \phi \quad (\Phi = 0)$$



FROM QUANTUM TO CLASSICAL

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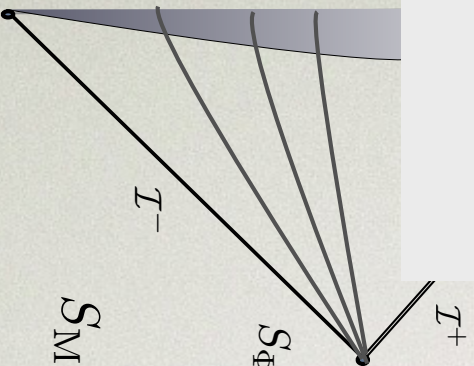
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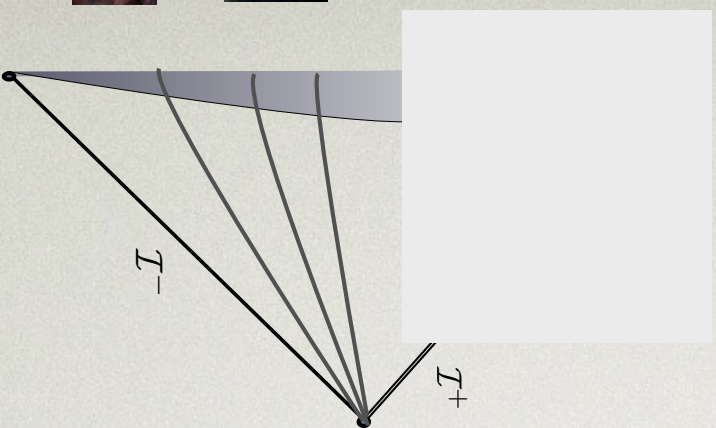
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Action describing the perturbations and its interaction

$$S_{\phi}[\phi] + S_{\gamma}[\gamma_{ab}] + S_{\text{int}}[\phi, \gamma_{ab}]$$



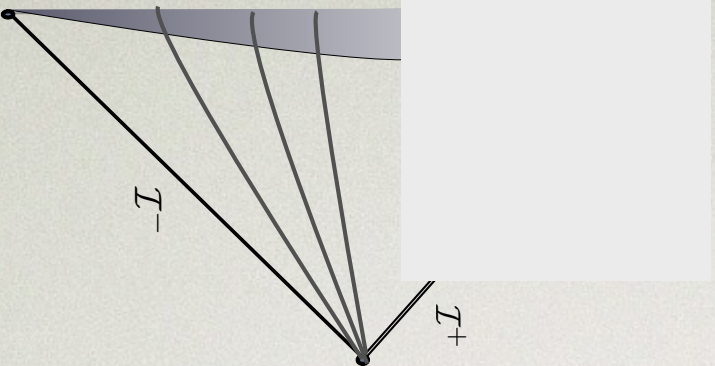
FROM QUANTUM TO CLASSICAL



FROM QUANTUM TO CLASSICAL

free scalar field

$$S_\phi[\phi] \equiv -\frac{1}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} \phi P \phi, \quad P \equiv -(\nabla^a \nabla_a + m^2 + \xi R)$$



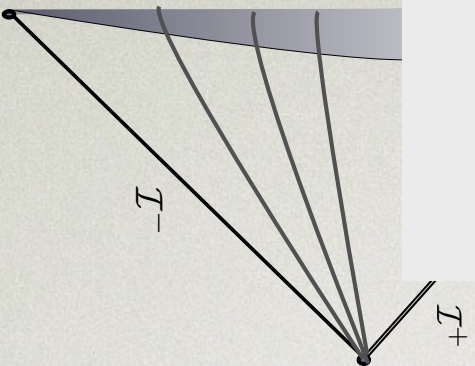
FROM QUANTUM TO CLASSICAL

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free graviton (with gauge fixing term)

$$S_\gamma[\gamma_{ab}] \equiv -\frac{1}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} \gamma^{ab} \mathfrak{D}_{abcd} \gamma^{cd}$$



FROM QUANTUM TO CLASSICAL

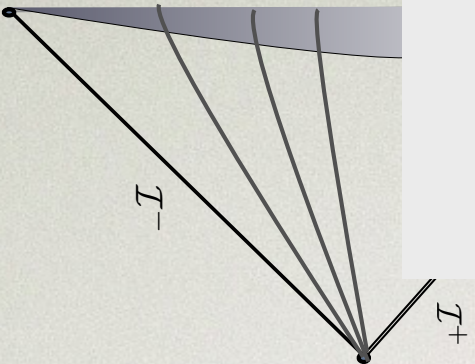
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$$\hat{\gamma}_{ab}(t, \mathbf{x}) = \sum_j \int \frac{d\mathfrak{y}(\alpha)}{\sqrt{2\omega_\alpha}} \hat{b}_\alpha^{(j)} e^{-i\omega_\alpha t} \varepsilon_{\alpha ab}^{(j)}(\mathbf{x}) + \text{H.c.}$$



FROM QUANTUM TO CLASSICAL

free scalar field

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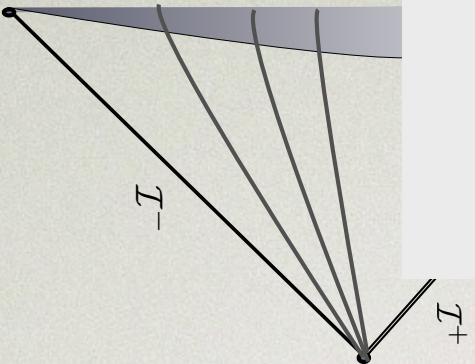
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Interaction

$$S_{\text{int}}[\phi, \gamma_{ab}] \equiv \int_{\mathcal{M}} d^4x \sqrt{-g} [\mathcal{L}_{\Phi}^{(3)} + \mathcal{L}_{\Phi}^{(4)}] \quad \mathcal{L}_{\Phi}^{(3)} \equiv \frac{\kappa}{2} T_{ab} \gamma^{ab}$$



FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



$$\phi = \phi_s + \phi_u$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



$$\phi = \phi_s + \phi_u$$

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

ϕ_u

System of interest - Unstable mode (which is the one we expect that "classicalize")

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



$$\phi = \phi_s + \phi_u$$

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$$\phi_u$$

Environment - Stable modes and graviton

$$\phi_s, \gamma_{ab}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



$$\phi = \phi_s + \phi_u$$

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$$\mathcal{S}_\phi[\phi] = \mathcal{S}_\phi[\phi_s] + \mathcal{S}_\phi[\phi_u]$$

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FROM QUANTUM TO CLASSICAL

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Σ_t



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$$\mathcal{S}_\phi[\phi] = \mathcal{S}_\phi[\phi_s] + \mathcal{S}_\phi[\phi_u]$$

Environment - Stable modes and graviton

$$\phi_s, \gamma_{ab}$$

$$T_{ab} = T_{ab}^{(s)} + T_{ab}^{(u)} + t_{ab}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



$$\phi = \phi_s + \phi_u$$

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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$$S_\phi[\phi] = S_\phi[\phi_s] + S_\phi[\phi_u]$$

Environment - Stable modes and graviton

$$\phi_s, \gamma_{ab}$$

$$T_{ab} = T_{ab}^{(s)} + T_{ab}^{(u)} + t_{ab}$$

$$S_{\text{int}}[\phi, \gamma_{ab}] \equiv \frac{\kappa}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} T_{ab}^{(s)} \gamma^{ab} + \frac{\kappa}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} T_{ab}^{(u)} \gamma^{ab} + \frac{\kappa}{2} \int_{\mathcal{M}} d^4x \sqrt{-g} t_{ab} \gamma^{ab}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



Initial state of the full system

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

$$\hat{\varrho}(0) = \hat{\varrho}_s(0) \otimes \hat{\varrho}_u(0) \otimes \hat{\varrho}_\gamma(0),$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



Initial state of the full system

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$$\hat{\varrho}_s(0) = |0_s\rangle\langle 0_s|$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



Initial state of the full system

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

$$\hat{\varrho}(0) = \hat{\varrho}_s(0) \otimes \hat{\varrho}_u(0) \otimes \hat{\varrho}_\gamma(0),$$

$$\hat{\varrho}_s(0) = |0_s\rangle\langle 0_s|$$

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FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

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$$\hat{\rho}_\gamma(0) = \text{thermal state with temperature } T$$

Initial state of the full system

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



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Initial state of the full system

$$\hat{\rho}(0) \xrightarrow{\hat{U}_{u\gamma s}(t)} \hat{\rho}(t)$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



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Initial state of the full system

$$\hat{\rho}(0) \xrightarrow[\text{trace out the environment}]{\hat{U}_{u\gamma s}(t)} \hat{\rho}(t)$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

Σ_t



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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Initial state of the full system

$$\hat{\rho}(0) \xrightarrow[\text{trace out the environment}]{\hat{U}_{u\gamma s}(t)} \hat{\rho}_u(t) = \text{tr}_{\gamma s} \hat{\rho}(t)$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



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Initial state of the full system

$$\hat{\rho}(0) \xrightarrow[\text{trace out the environment}]{\hat{U}_{u\gamma s}(t)} \hat{\rho}_u(t) = \text{tr}_{\gamma s} \hat{\rho}(t)$$

$$\varrho_{\text{red}}(t, \varphi_u, \varphi_u') = \int d\psi_u d\psi_u' \varrho_u(0, \psi_u, \psi_u') \times J_{\text{red}}(t, \varphi_u, \varphi_u'; 0, \psi_u, \psi_u'),$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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$$\hat{\varrho}_\gamma(0) = \text{thermal state with temperature } T$$

Initial state of the full system

$$\hat{\rho}(0) \xrightarrow[\text{trace out the environment}]{\hat{U}_{u\gamma s}(t)} \hat{\rho}_u(t)$$

$$\hat{\rho}_u(t) = \text{tr}_{\gamma s} \hat{\rho}(t)$$

$$\varrho_u(0, \psi_u, \psi_u') \equiv \langle \psi_u | \hat{\varrho}_u(0) | \psi_u' \rangle$$

$$\varrho_{\text{red}}(t, \varphi_u, \varphi_u') = \int d\psi_u d\psi_u' \varrho_u(0, \psi_u, \psi_u') \times J_{\text{red}}(t, \varphi_u, \varphi_u'; 0, \psi_u, \psi_u'),$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



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$$\hat{\rho}(0) \xrightarrow[\text{trace out the environment}]{\hat{U}_{u\gamma s}(t)} \hat{\rho}_u(t) = \text{tr}_{\gamma s} \hat{\rho}(t)$$

$$\varrho_{\text{red}}(t, \varphi_u, \varphi_u') = \int d\psi_u d\psi_u' \varrho_u(0, \psi_u, \psi_u') \times J_{\text{red}}(t, \varphi_u, \varphi_u'; 0, \psi_u, \psi_u'),$$

$$\hat{\phi}_u(0, \mathbf{x})|\psi_u\rangle = \psi_u(\mathbf{x})|\psi_u\rangle$$

$$\varrho_u(0, \psi_u, \psi_u') \equiv \langle \psi_u | \hat{\varrho}_u(0) | \psi_u' \rangle$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



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$$\varrho_u(0, \psi_u, \psi_u') \equiv \langle \psi_u | \hat{\varrho}_u(0) | \psi_u' \rangle$$

$$\hat{\phi}_u(0, \mathbf{x}) | \psi_u \rangle = \psi_u(\mathbf{x}) | \psi_u \rangle$$

$$J_{\text{red}}(t, \varphi_u, \varphi_u'; 0, \psi_u, \psi_u') \equiv \int_{\psi_u}^{\varphi_u} \mathcal{D}\phi_u \int_{\psi_u'}^{\varphi_u'} \mathcal{D}\phi_u' e^{i\{S_\phi[\phi_u] - S_\phi[\phi_u']\}} F[\phi_u, \phi_u']$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

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$$J_{\text{red}}(t, \varphi_u, \varphi_u'; 0, \psi_u, \psi_u') \equiv \int_{\psi_u}^{\varphi_u} \mathcal{D}\phi_u \int_{\psi_u'}^{\varphi_u'} \mathcal{D}\phi_u' e^{i\{S_\phi[\phi_u] - S_\phi[\phi_u']\}} F[\phi_u, \phi_u']$$

Feynman-Vernon Influence Functional
 $\varrho_u(0, \mathbf{x})|\psi_u\rangle = |\psi_u(\mathbf{x})|\psi_u\rangle$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

Σ_t



$$F[\phi_u, \phi_u'] \equiv \int d\varphi_s \int d\psi_s d\psi_s' \varrho_s(0, \psi_s, \psi_s') \int_{\psi_s, \psi_s'}^{\varphi_s = \varphi_s'} \mathcal{D}\phi_s \mathcal{D}\phi_s' e^{i\{S_\phi[\phi_s] - S_\phi[\phi_s']\}} \tilde{F}[\phi_s + \phi_u, \phi_s' + \phi_u'],$$

$$\tilde{F}[\phi, \phi'] \equiv \int d\varsigma_{ab} \int d\xi_{ab} d\xi_{ab}' \varrho_\gamma(0, \xi_{ab}, \xi_{ab}') \int_{\xi_{ab}, \xi_{ab}'}^{\varsigma_{ab} = \varsigma_{ab}'} \mathcal{D}\gamma_{ab} \mathcal{D}\gamma_{ab}' e^{i\{S_\gamma[\gamma_{ab}] - S_\gamma[\gamma_{ab}']\}} e^{i\{S_{\text{int}}[\phi, \gamma_{ab}] - S_{\text{int}}[\phi', \gamma_{ab}']\}}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



**Master equation for the
state of the unstable mode**

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

Σ_t



Master equation for the state of the unstable mode

$$\begin{aligned} \partial_t \varrho_{\text{red}} = & -i \left[\frac{1}{2} (-\partial_q^2 + \partial_{q'}^2) - \frac{\Omega^2}{2} (q^2 - q'^2) \right] \varrho_{\text{red}} - \frac{\kappa^2}{4} \int_0^t d\tau [\text{Re} \langle \hat{Q}(\tau) \hat{Q}(0) \rangle \cosh \Omega \tau - \Omega \text{Re} \langle \hat{Q}(\tau) \hat{P}(0) \rangle \sinh \Omega \tau] (q - q')^2 \varrho_{\text{red}} \\ & - \frac{\kappa^2}{4} \int_0^t d\tau [\text{Re} \langle \hat{P}(\tau) \hat{P}(0) \rangle \cosh \Omega \tau - \Omega^{-1} \text{Re} \langle \hat{P}(\tau) \hat{Q}(0) \rangle \sinh \Omega \tau] (-i\partial_q - i\partial_{q'})^2 \varrho_{\text{red}} \\ & + \frac{\kappa^2}{4} \int_0^t d\tau \{ [\Omega^{-1} \text{Re} \langle \hat{Q}(\tau) \hat{Q}(0) \rangle + \Omega \text{Re} \langle \hat{P}(\tau) \hat{P}(0) \rangle] \sinh \Omega \tau \\ & - [\text{Re} \langle \hat{P}(\tau) \hat{Q}(0) \rangle + \text{Re} \langle \hat{Q}(\tau) \hat{P}(0) \rangle] \cosh \Omega \tau \} (q - q') (-i\partial_q - i\partial_{q'}) \varrho_{\text{red}} + \dots \end{aligned}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



**Wigner function for the
unstable mode**

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

Σ_t



Wigner function for the unstable mode

$$\begin{aligned} \partial_t W_{\text{red}} = & \{H(q, p), W_{\text{red}}\} + \frac{\kappa^2}{4} \int_0^t d\tau [\text{Re}\langle \hat{Q}(\tau) \hat{Q}(0) \rangle \cosh \Omega \tau - \Omega \text{Re}\langle \hat{Q}(\tau) \hat{P}(0) \rangle \sinh \Omega \tau] \partial_p^2 W_{\text{red}} \\ & + \frac{\kappa^2}{4} \int_0^t d\tau [\text{Re}\langle \hat{P}(\tau) \hat{P}(0) \rangle \cosh \Omega \tau - \Omega^{-1} \text{Re}\langle \hat{P}(\tau) \hat{Q}(0) \rangle \sinh \Omega \tau] \partial_q^2 W_{\text{red}} \\ & + \frac{\kappa^2}{4} \int_0^t d\tau \{ [\Omega^{-1} \text{Re}\langle \hat{Q}(\tau) \hat{Q}(0) \rangle + \Omega \text{Re}\langle \hat{P}(\tau) \hat{P}(0) \rangle] \sinh \Omega \tau \\ & - [\text{Re}\langle \hat{P}(\tau) \hat{Q}(0) \rangle + \text{Re}\langle \hat{Q}(\tau) \hat{P}(0) \rangle] \cosh \Omega \tau \} \partial_p \partial_q W_{\text{red}} + \dots \end{aligned}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

Σ_t



$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Long-time regime and high temperature limit

$$\Omega t \gg 1 \quad k_B T \gg \Omega$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



Master equation for the state of the unstable mode

$$\partial_t \varrho_{\text{red}} \approx -i \left[\frac{1}{2} (-\partial_q^2 + \partial_{q'}^2) - \frac{\Omega^2}{2} (q^2 - q'^2) \right] \varrho_{\text{red}} - D(q - q')^2 \varrho_{\text{red}} + \dots$$

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Wigner function for the unstable mode

$$\partial_t W_{\text{red}} \approx \{H(q, p), W_{\text{red}}\} + D \partial_p^2 W_{\text{red}} + \dots$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

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$$D = 8\pi \Delta \left(\frac{\ell_P}{R} \right)^2 \left(\frac{k_B T}{\Omega} \right) \Omega^2$$

FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



**Wigner function for the
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- ❖ L. Diósi and C. Kiefer J.Phys.A (2002)
- ❖ O. Brodier and A. Ozorio de Almeida, PRE (2004)

FROM QUANTUM TO CLASSICAL

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$$W_{\text{red}}(t > t_d) > 0 \text{ (regardless of the initial state)}$$

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- ❖ For a $R \sim 10 \text{ km}$ neutron star and $T \sim 1 \text{ K}$

FROM QUANTUM TO CLASSICAL

- Unstable Phase

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- For a $R \sim 10 \text{ km}$ neutron star and $T \sim 1 \text{ K}$

$$t_d \sim 160 \times \Omega^{-1} \sim 160 \times R$$

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FROM QUANTUM TO CLASSICAL

* Unstable Phase

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* For a $R \sim 10 \text{ km}$ neutron star and $T \sim 1 \text{ K}$

$$t_d \sim 160 \times \Omega^{-1} \sim 160 \times R \quad \text{which is of the order of} \quad t_{\text{br}} \sim 10^{-3} \text{ s}$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

Σ_t



FROM QUANTUM TO CLASSICAL

- ❖ Unstable Phase

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$$\varrho_{\text{red}}(t, q, q') \approx 2e^{-2\sigma^2(q-q')^2 - i\Omega(q^2 - q'^2)/2} e^{-\Omega t} \int_{-\infty}^{+\infty} du' \tilde{W}_{\text{red}}(0, u', 0)$$

FROM QUANTUM TO CLASSICAL

❖ Unstable Phase

$$ds^2 = -f(dt^2 - d\chi^2) + r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

Σ_t



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Pointer states are narrow enough (in the long-time regime) to be approximated by classical states

FROM QUANTUM TO CLASSICAL

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Σ_t



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mixture of localized states with width:

THUS, BY THE TIME BACKREACTION BECOMES INELUCTABLE THE (UNSTABLE SECTOR OF THE) INITIALLY PURE VACUUM STATE HAS EVOLVED INTO A STATISTICAL MIXTURE OF LOCALIZED STATES IN THE FIELD AMPLITUDE AND MOMENTUM REPRESENTATION.

$$(2\sigma)^{-1} \sim \frac{\chi_P}{R} \sqrt{k_B T}$$

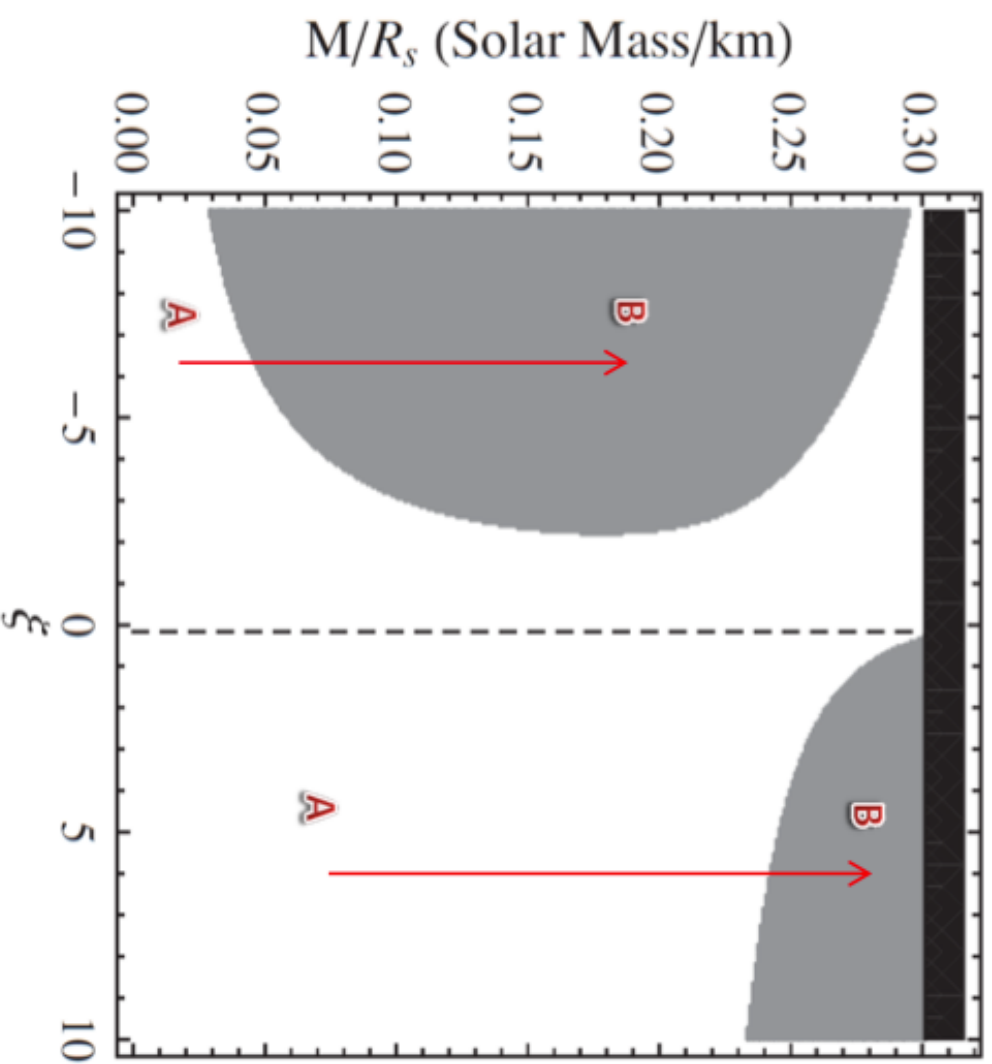
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BACKREACTION



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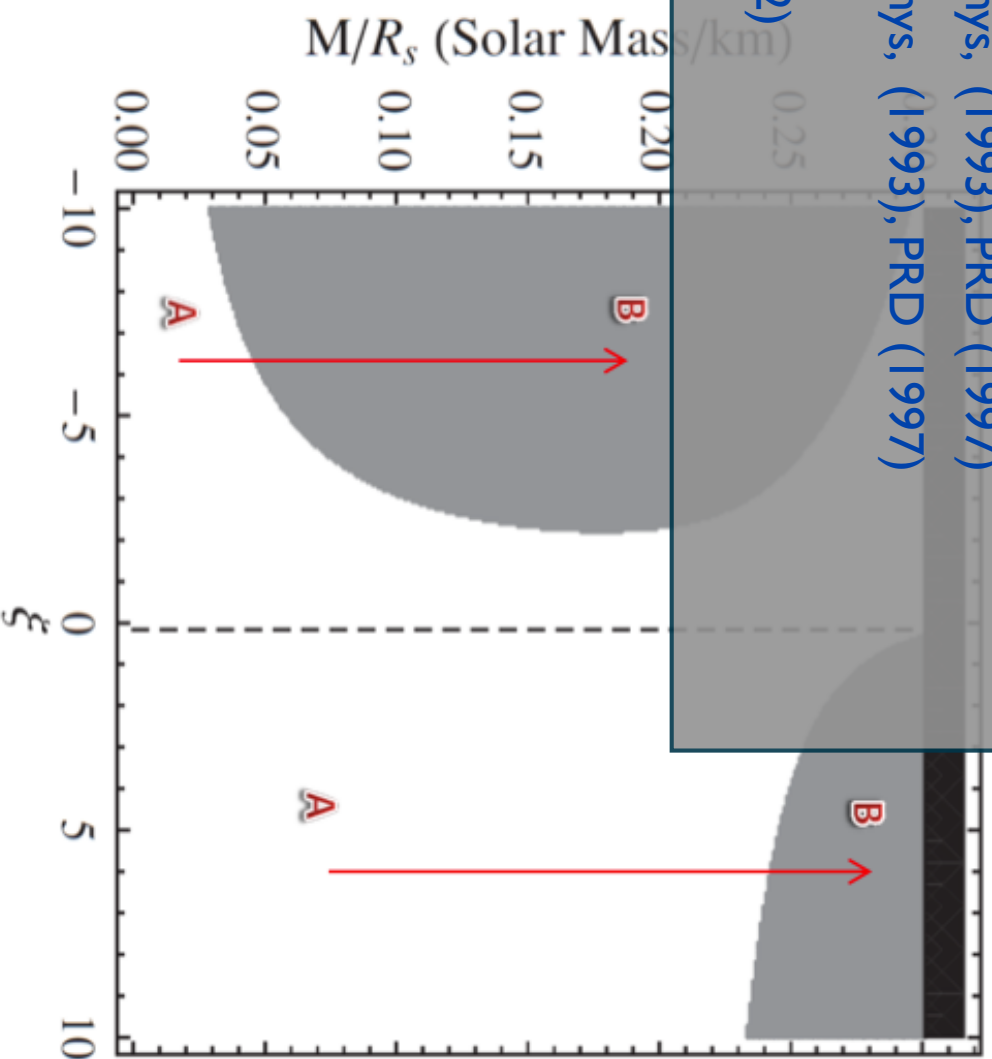
T. Damour, E. Esposito-Farèse, PRL (1993), PRD (1996)

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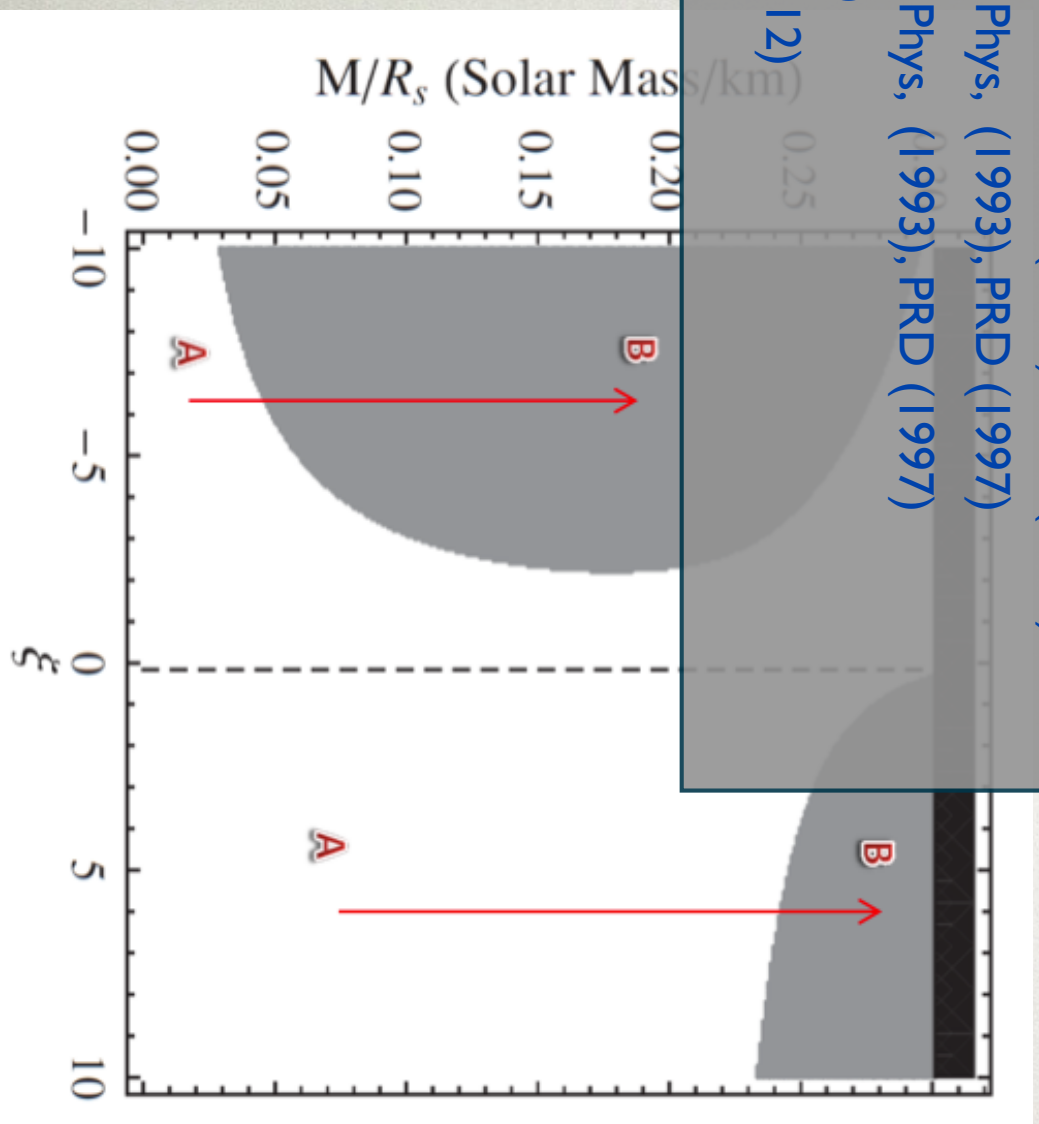
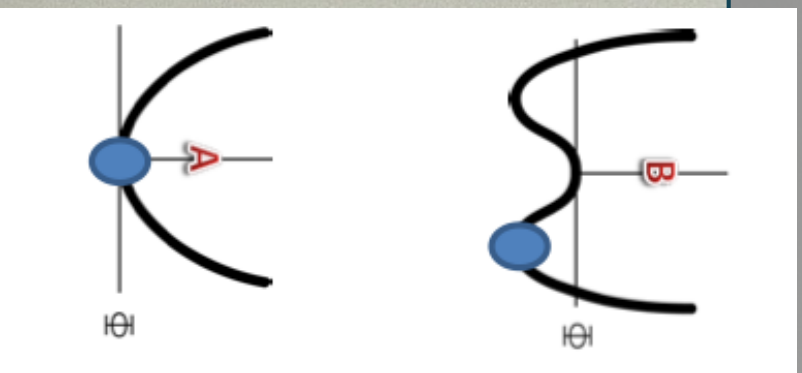
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